

Magnetized plasma plumes: physics, open problems

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EXB Plasmas for Space and Industrial Applications

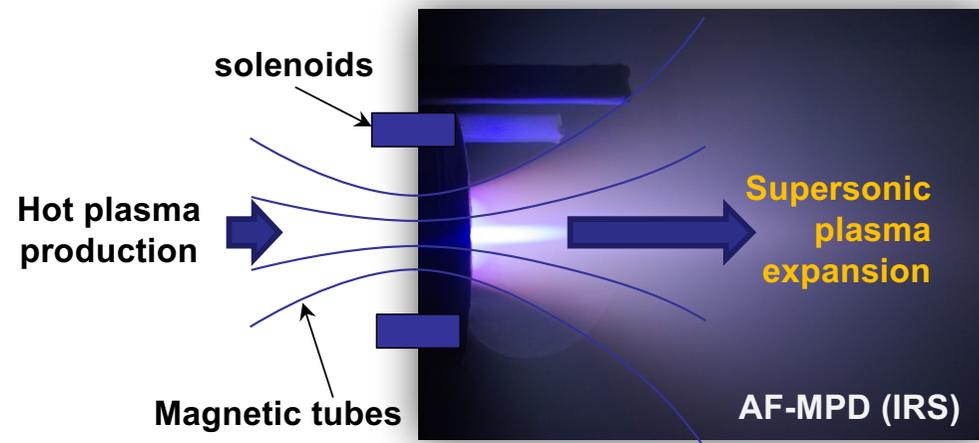
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Plumes generated by magnetic nozzles

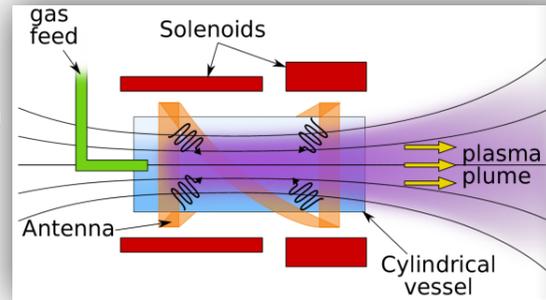
- A magnetic nozzle (MN) is the convergent-divergent magnetic topology created by an axisymmetric set of coils or permanent magnets,
- A MN is supposed to:
 - channel and accelerate a plasma plume
 - reduce plasma backflow
 - produce (magnetic) thrustin a way similar to a solid deLaval nozzle does with a hot, neutral gas.
- Advantages:
 - MN has no walls → no heat loads, no erosion
 - MN shape is adaptable → thruster throttling (I_{sp} vs thrust)
 - 3D MN → thrust vector control
- The MN is being proposed as the main acceleration stage of several new electromagnetic thrusters featuring quasi-axial magnetic topologies.



Plasma thrusters with MN (I)

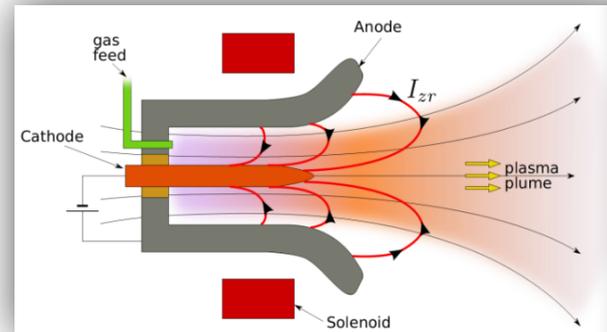
- **The Helicon Plasma Thruster (HPT)**

- High plasma density ($10^{18} - 10^{20} \text{ m}^{-3}$)
- Prototypes in the 50 W – 50 kW range, but still low efficiency: $\eta_T \ll 0.5$



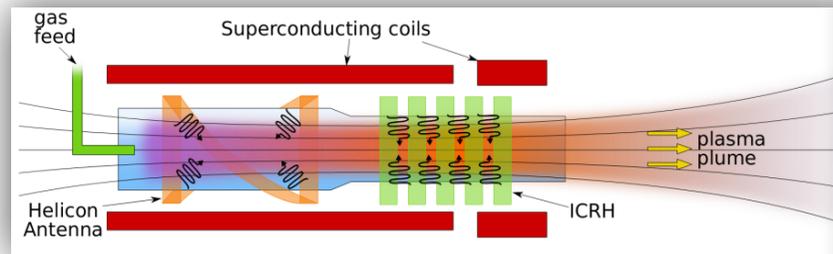
- **AF-MPD**

- DC discharge
- High power (10 kW – 200 kW); $\eta_T < 0.5$



- **VASIMR**

- A HPT enhanced with an ion-cyclotron-resonance heater stage
- Requires much higher magnetic field
- High power (100 kW – few MW) $\eta_T < 0.72$



HELICON (GEORGIA U.)



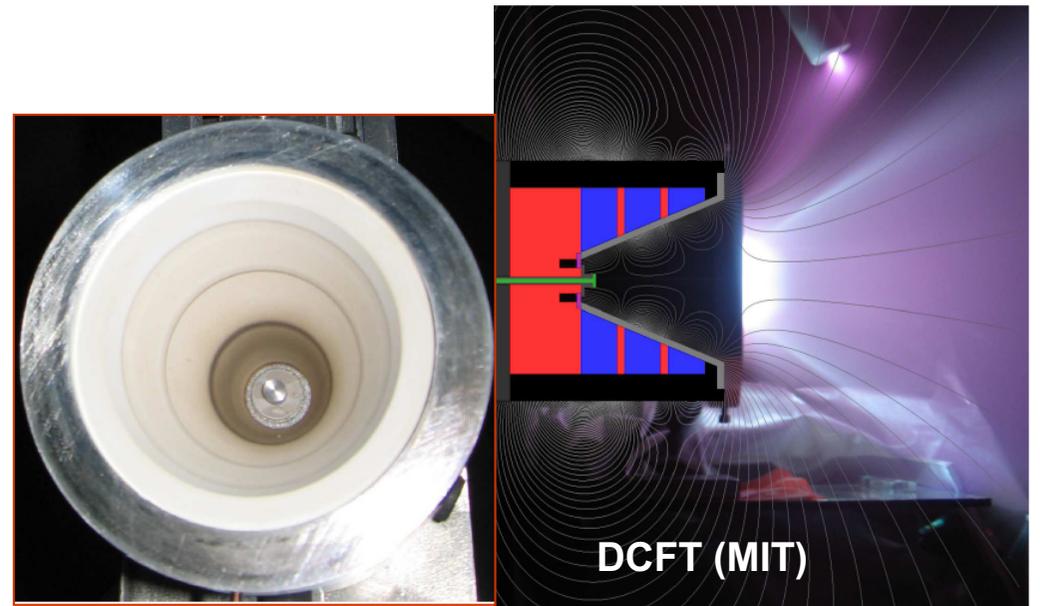
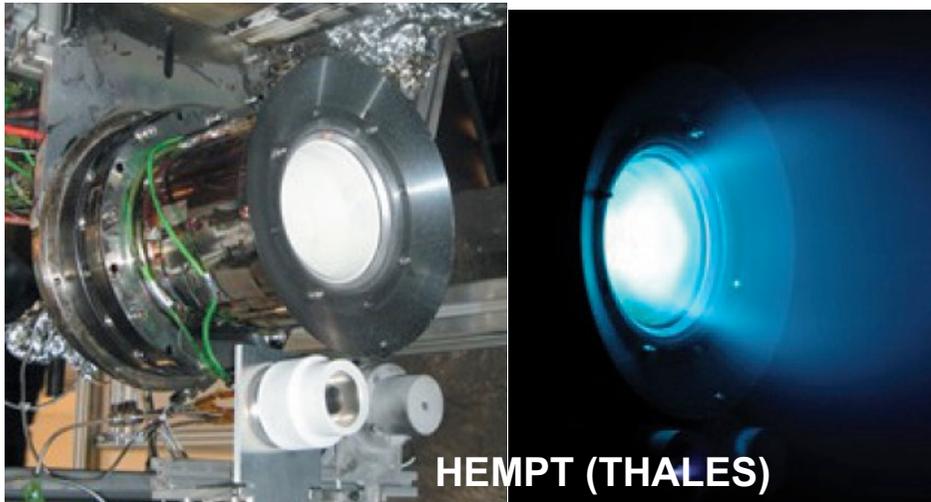
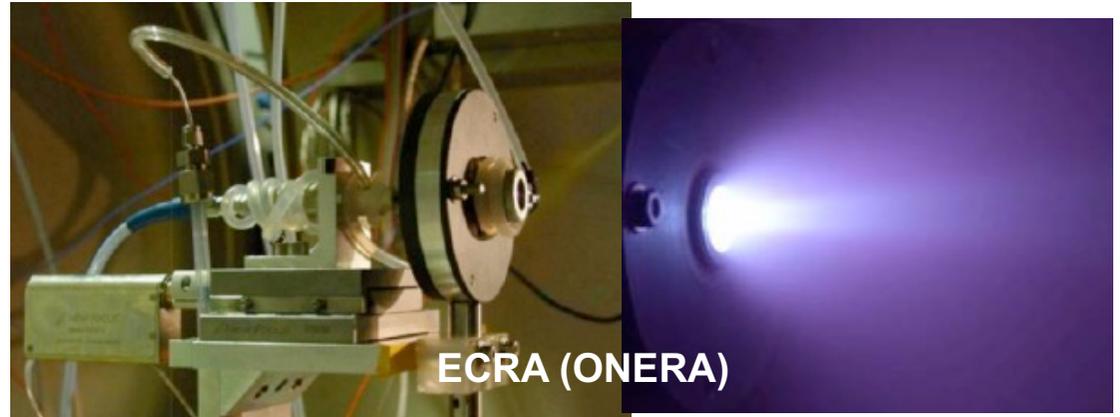
AF-MPD (STUTTGART U.)



VASIMR (ASTRA)

Plasma thrusters with MN (II)

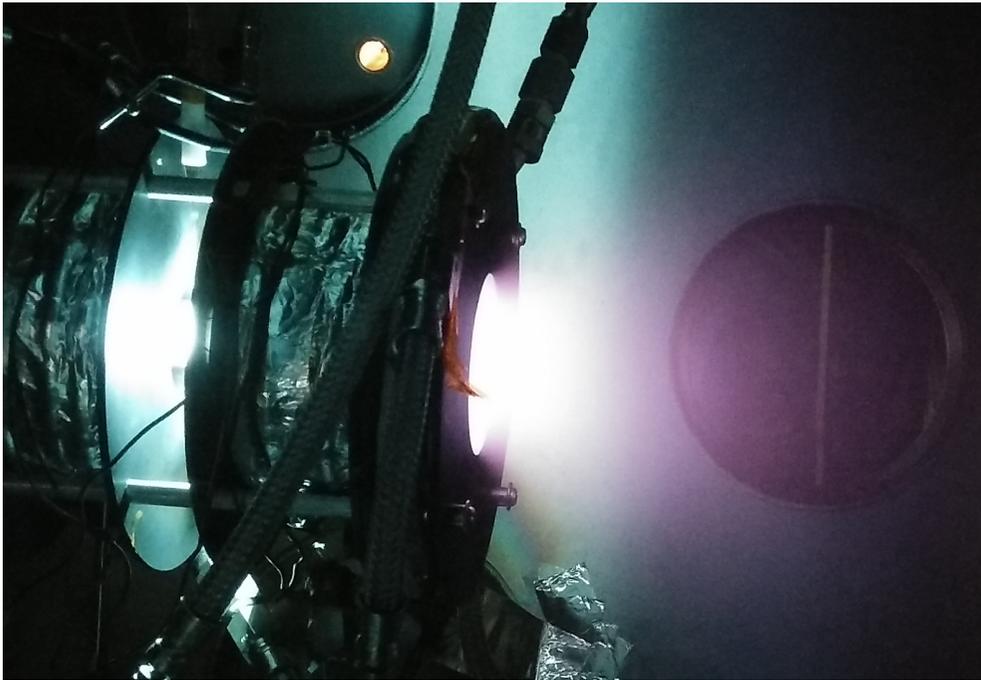
- **The Electron Cyclotron Resonance Thruster**
 - Similar to HPT but different plasma-wave interaction
- All these thrusters share **no external neutralizer**
- All, except AFMPDT, have no electrodes
- **HEMPT** and **DCFT** (variants of Hall thrusters) include a MN but this should have a minor role, ,since the plasma is already expanded



MN channelling effect

- HPT05 prototype (EP2-SENER) with the MN-coil turned

OFF

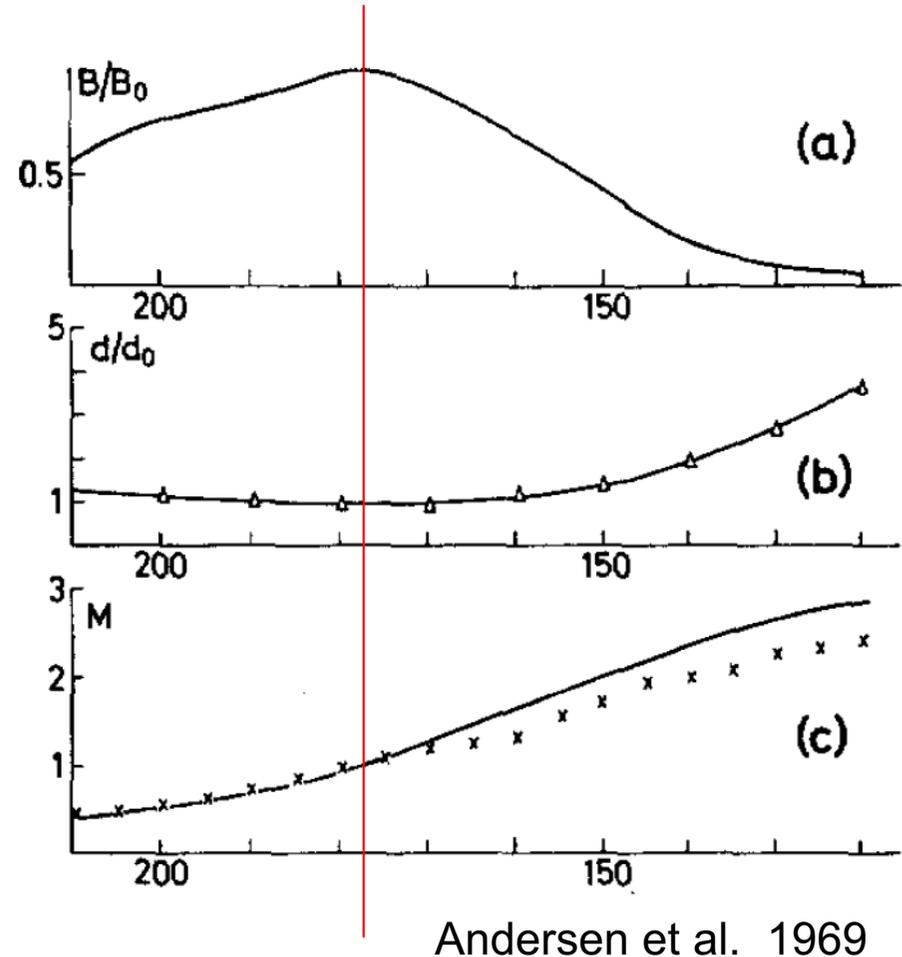


ON



When the MN principle was proven?

- The magnetic/solid nozzle analogy was demonstrated experimentally in 1969
- A few prototypes with MNs were studied in the 1970s-1980s.
- The emergence of research on new thrusters in the last decade has boosted the interest in MN physics and performances.
- Also, MN coupled to plasma sources are being used for plasma manufacturing processes and supersonic wind tunnels



What do we need to understand of MN physics?

Study is focused on **MN divergent region**, outside the thruster chamber

- Convergent region is inside chamber usually, and magnetic confinement there is strongly coupled with chamber processes (ionization, heating, wall-interaction,...)
- **2D supersonic expansion** features
 - Can a quasi-1D model describe MN physics? (Very limited way)
- Effectiveness of **magnetic confinement**
 - Must ions be magnetized?
- Plasma **energy conversion** process from 'internal' to 'axial kinetic' one
- Influence of the **type of internal energy** in the above process (electron/ion, isotropic/ anisotropic, ...) **different plasma sources** → **different MNs**
- Downwards **detachment** of plasma from MN
- **Thrust transmission** mechanism:
 - different from solid nozzle (based on pressure on oblique walls) !
- Others, less obvious:
 - collisionless plume cooling,
 - induced B-field,...

Main assumptions for core MN model

- Plasma conditions at MN throat known; R = MN radius at throat

- Quasineutral, current-free plasma

$$\lambda_{d0}/R \rightarrow 0, \quad n_e = n_i \equiv n$$

- Collisionless, fully-ionized plasma $\Omega_{e0}/\nu_{e0} \rightarrow \infty$,

- Relevant for propulsive application

- Two-fluid model

- Electron inertia is neglected $m_e/m_i \rightarrow 0$

- Fully-magnetized electrons, partially-magnetized ions

$$\ell_{e0}/R \rightarrow 0 \quad \ell_{i0}/R = O(1)$$

- Effects of induced magnetic field on plasma flow are negligible

- Moderate plasma density

$$\beta_0 \rightarrow 0, \text{ with } \beta = \mu_0 n T_e / B^2$$

- Internal energy is stored isotropically in electrons; $T_i \ll T_e$

- Application limited to HPT and partially to ECRT

DIMAGNO: a 2D fluid formulation

- Fluid equations

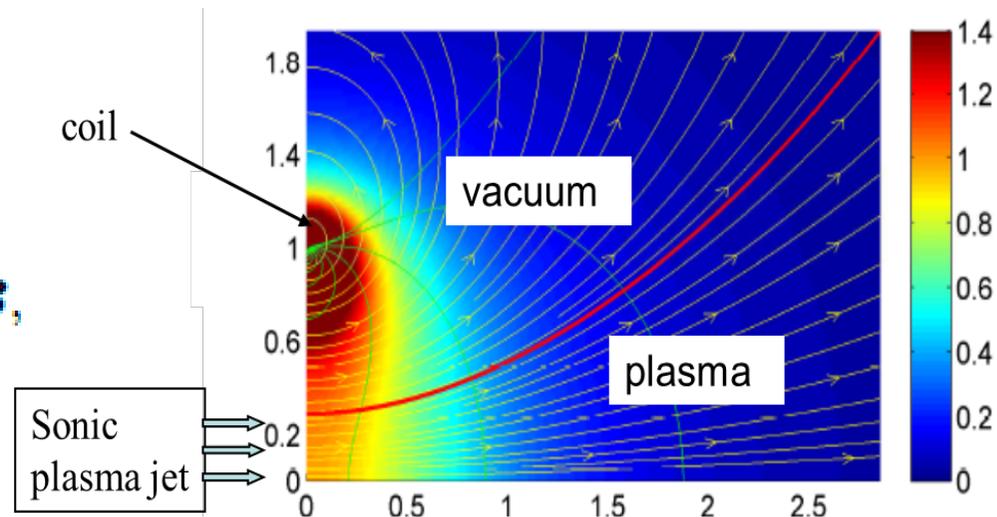
$$\nabla \cdot (n\mathbf{u}_i) = 0,$$

$$\nabla \cdot (n\mathbf{u}_e) = 0,$$

$$m_i n (\mathbf{u}_i \cdot \nabla) \mathbf{u}_i = -en\nabla\phi + en\mathbf{u}_i \times \mathbf{B},$$

$$0 = -\nabla p_e + en\nabla\phi - en\mathbf{u}_e \times \mathbf{B},$$

$$\frac{T_e}{T_0} = \left(\frac{n}{n_0}\right)^{\gamma-1}$$

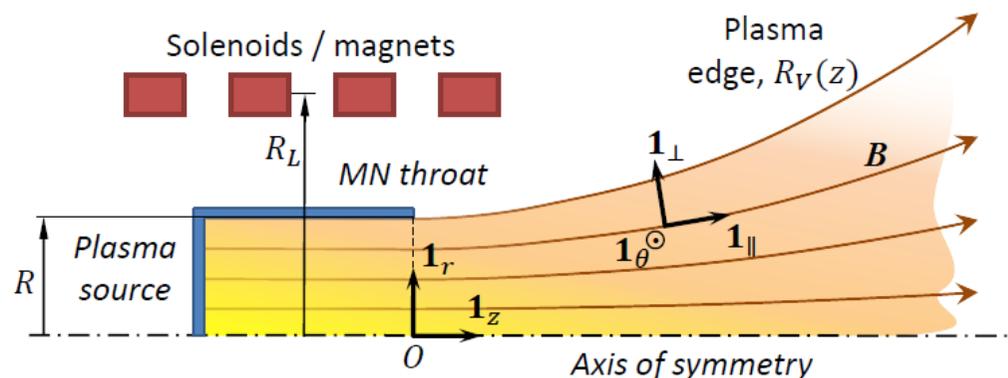


- 9 (scalar) equations for 9 variables

- 2 systems of reference

$$\{\mathbf{1}_z, \mathbf{1}_r, \mathbf{1}_\theta\} \quad \{\mathbf{1}_\parallel, \mathbf{1}_\perp, \mathbf{1}_\theta\}$$

- Streamtubes for
 - B-field
 - Ion fluid
 - electron fluid



DIMAGNO: a 2D fluid formulation

- Magnetic streamfunction for solenoidal longitudinal B field

$$\nabla\psi = rB\mathbf{1}_\perp : \partial\psi/\partial z = -rB_r, \quad \partial\psi/\partial r = rB_z,$$

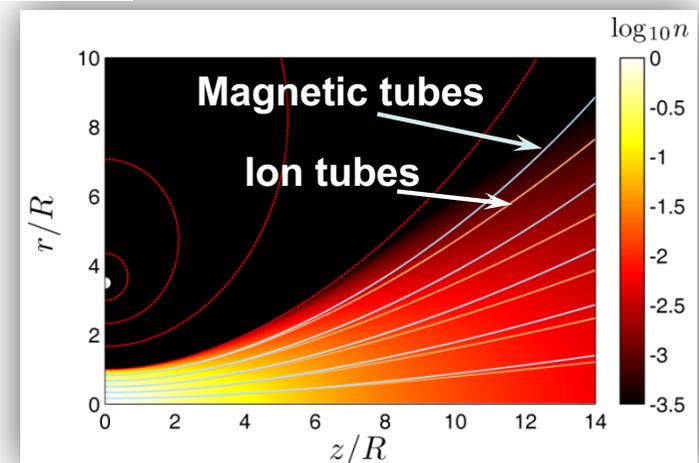
- B- streamtubes (constant magnetic flux) $\psi(r, z) = \text{const}$

- Plasma jet edge corresponds to $\psi_V(z, r) = \psi(0, R)$

- Ion/electron streamfunctions

$$\frac{\partial\psi_j}{\partial z} = -rn u_{rj} \quad \frac{\partial\psi_j}{\partial r} = rn u_{zj}, \quad j = i, e$$

- i,e - streamtubes: $\psi_i, \psi_e = \text{const}$:
particle flux of ions/electrons = const



- Electric current density: $\mathbf{j} = en(\mathbf{u}_i - \mathbf{u}_e)$

- It is convenient to split vector variables ($\mathbf{j}, \mathbf{B}, \mathbf{u}_i, \mathbf{u}_e, \dots$) in

- Azimuthal: $j_\theta = \mathbf{j} \cdot \mathbf{1}_\theta$
- Longitudinal: $\tilde{\mathbf{j}} = \mathbf{j} - j_\theta \mathbf{1}_\theta$

MN conditions at the throat

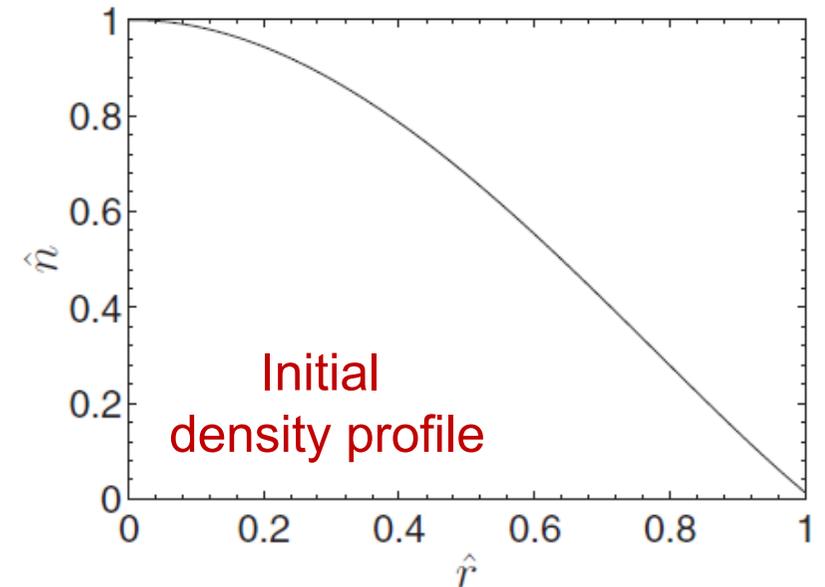
- Conditions at MN throat are initial integration conditions
- They are representative of plasma source characteristics
- Here we apply conditions derived from models of a cylindrical HPT source (with strong wall magnetic shielding)

$$n(0, r) \text{ known}, \quad \phi(0, r) \approx 0,$$

$$u_{\theta i}(0, r) \approx 0, \quad u_{\theta e}(0, r) = - \left[\frac{\partial p_e / \partial r}{enB} \right]_{z=0},$$

$$u_{ri}(0, r) = u_{re}(0, r) \approx 0,$$

$$u_{zi}(0, r) = M_0 c_{s0}, \quad (M_0 \geq 1), \quad \int_0^R n u_{ze} r dr = \int_0^R n u_{zi} r dr$$



Conservation/algebraic relations for electrons

- Electron equations are reduced to simple conservation equations

$$\begin{aligned}
 u_{\perp e} &= 0 \\
 nu_{\parallel e}/B &= G_e(\psi), \\
 h_e(n) - e\phi &= H_e(\psi). \\
 u_{\theta e} &= -\frac{E_{\perp}}{B} - \frac{1}{enB} \frac{\partial p_e}{\partial \mathbf{1}_{\perp}} \equiv -\frac{r}{e} \frac{dH_e}{d\psi},
 \end{aligned}$$

$$\begin{aligned}
 h_e(n) &= T_{e0} \ln n, & \text{for } \gamma_e = 1, \\
 h_e(n) &= T_e \frac{\gamma}{\gamma - 1}, & \text{for } \gamma > 1.
 \end{aligned}$$

- Functions G_e and H_e are defined at the MN throat

- 1st eq. : electrons remain in the same B-streamsurface
- 2nd eq. : e-streamtubes = B-streamtubes, i.e. $\psi_e(\psi)$
- 3rd eq. : General Boltzmann equilibrium applies in each e-streamsurface
[e.g. for $T_e = \text{const}$: $n \propto \exp(e\phi/T_{e0})$]
- 4th eq. : Azimuthal e-drift is the combination of $E \times B$ and $\nabla p_e \times B$ drifts
 - (Expanding) pressure \perp -force = (confining) electric + magnetic \perp - forces
 - Electron fluid iso-rotates ($\frac{u_{\theta e}}{r} = \text{const}$) in its helicoidal drift within B-streamsurface

Partial differential equations for ions

- Three hyperbolic differential equations for n , u_{zi} and u_{ri}

$$u_{ri} \frac{\partial \ln n}{\partial r} + u_{zi} \frac{\partial \ln n}{\partial z} + \frac{\partial u_{ri}}{\partial r} + \frac{\partial u_{zi}}{\partial z} = -\frac{u_{ri}}{r},$$

$$u_{ri} \frac{\partial u_{ri}}{\partial r} + u_{zi} \frac{\partial u_{ri}}{\partial z} + c_s^2 \frac{\partial \ln n}{\partial r} = -(u_{\theta e} - u_{\theta i}) \Omega_i \cos \alpha + \frac{u_{\theta i}^2}{r},$$

$$u_{ri} \frac{\partial u_{zi}}{\partial r} + u_{zi} \frac{\partial u_{zi}}{\partial z} + c_s^2 \frac{\partial \ln n}{\partial z} = (u_{\theta e} - u_{\theta i}) \Omega_i \sin \alpha,$$

Ion thermal speed:

$$c_s = \sqrt{\gamma T_e / m_i}$$

- plus conservation of azimuthal canonical momentum for $u_{\theta i}$

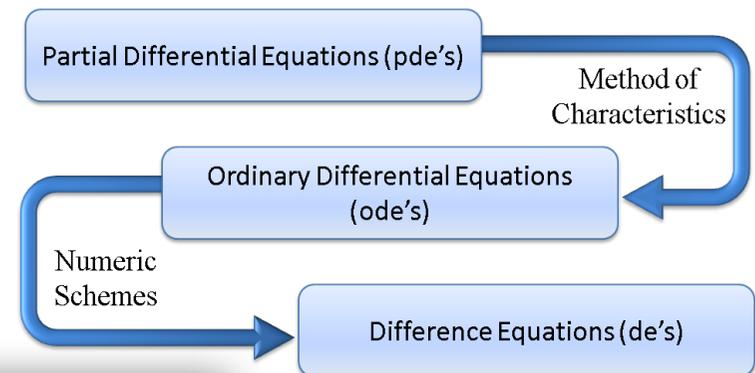
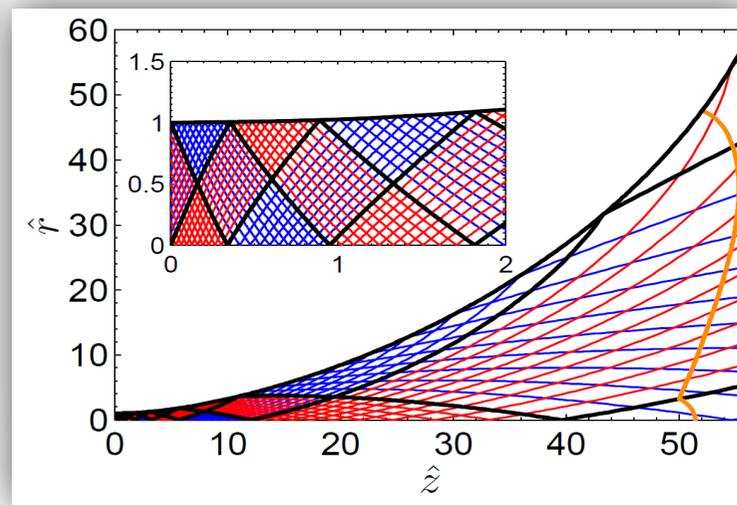
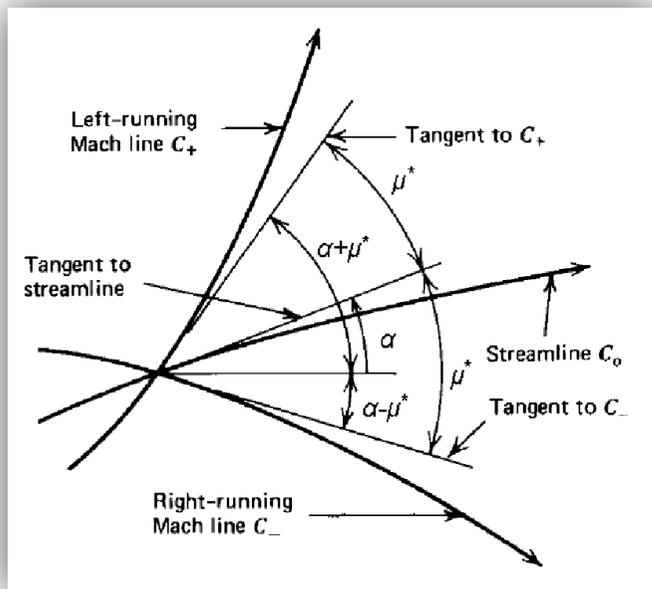
$$m_i (\tilde{\mathbf{u}}_i \cdot \nabla) r u_{\theta i} = -r e u_{\perp i} B = -e (\tilde{\mathbf{u}}_i \cdot \nabla) \psi$$

$$\longrightarrow r m_i u_{\theta i} + e \psi = D_i(\psi_i) \quad (D_i \text{ defined at the MN throat})$$

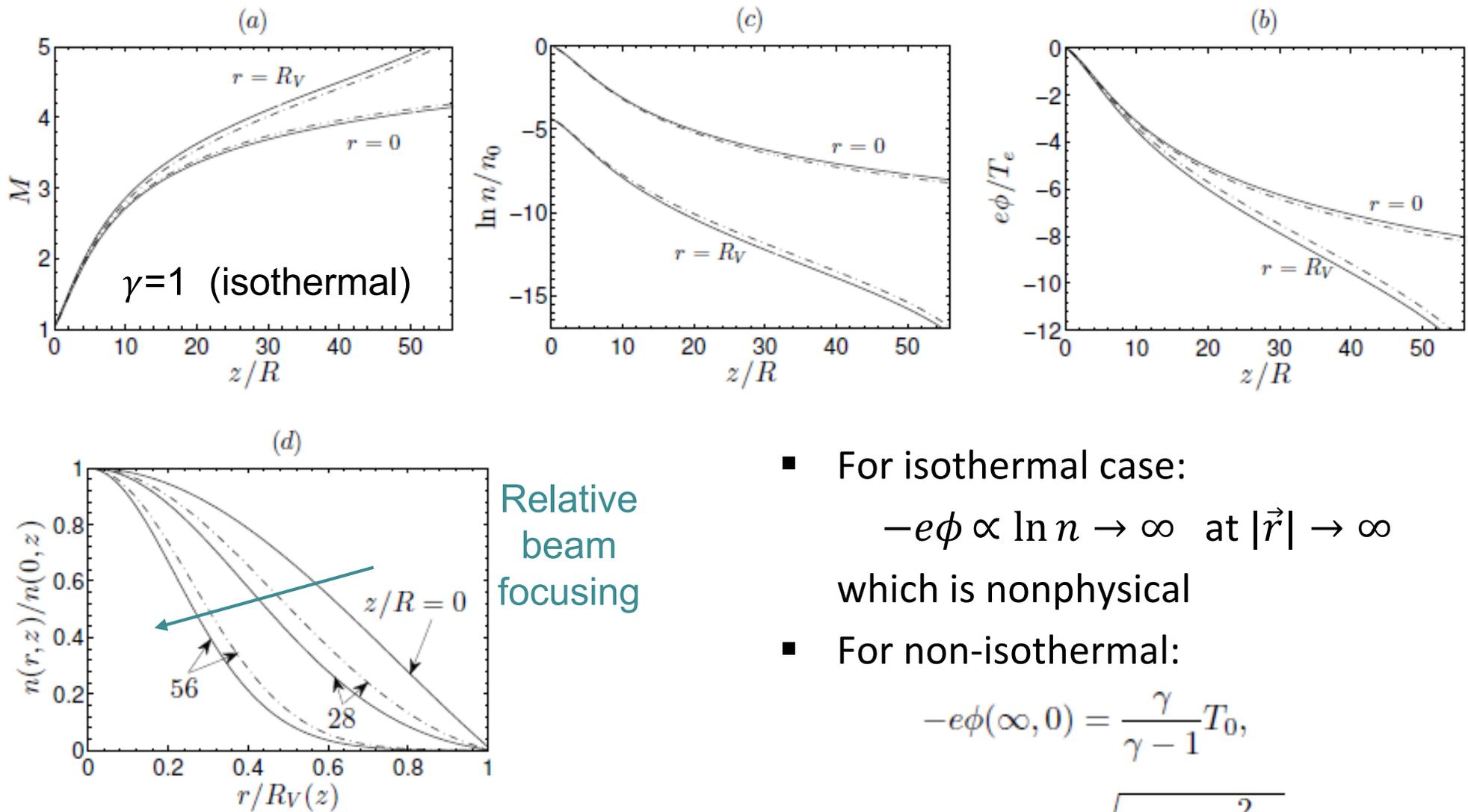
- Mechanisms governing azimuthal velocities are different for inertial ions and confining electrons

Integrating with method of characteristics

- Mach number for longitudinal ion velocity: $M = \tilde{u}/c_s$
- If the longitudinal ion velocity is supersonic ($M > 1$), the above PDEs can be integrated forward-marching with method of characteristic lines.
 - At each point, these are 3: the ion-streamline + 2 Mach-lines



The 2D plasma jet expansion



- For isothermal case:
 - $-e\phi \propto \ln n \rightarrow \infty$ at $|\vec{r}| \rightarrow \infty$ which is nonphysical
- For non-isothermal:

$$-e\phi(\infty, 0) = \frac{\gamma}{\gamma - 1} T_0,$$

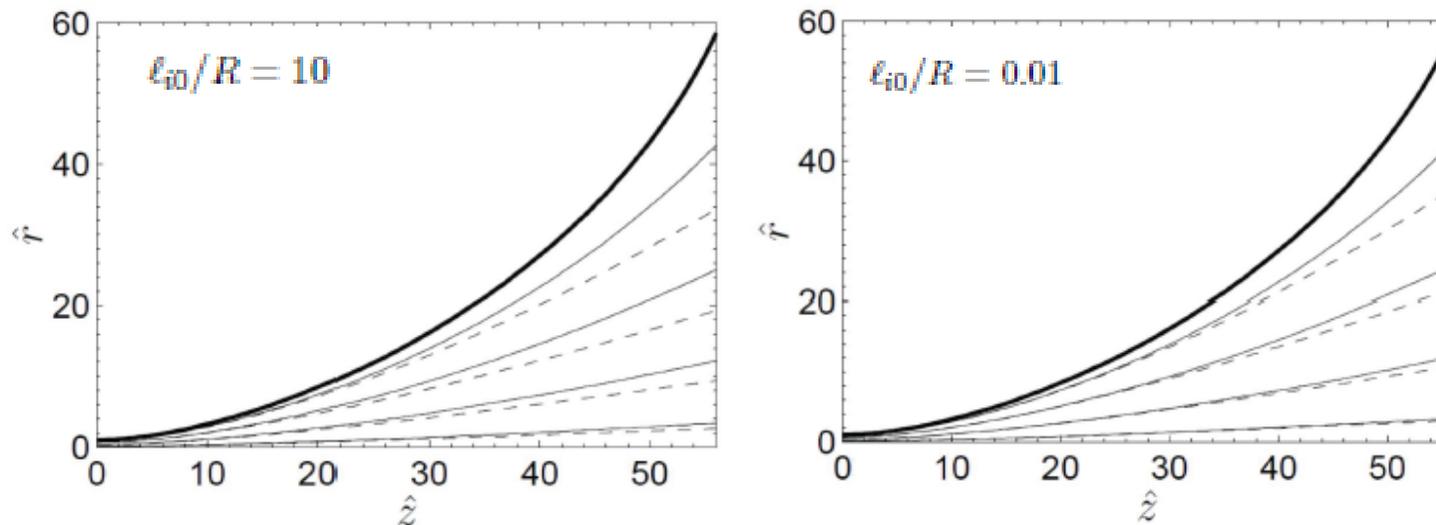
$$u_i(\infty, 0) = c_{s0} \sqrt{M_0^2 + \frac{2}{\gamma - 1}},$$

Ion separation

- Electron streamlines are bended radially into B-lines by full-magnetization
- Ion streamlines are bended radially mainly by the electric field

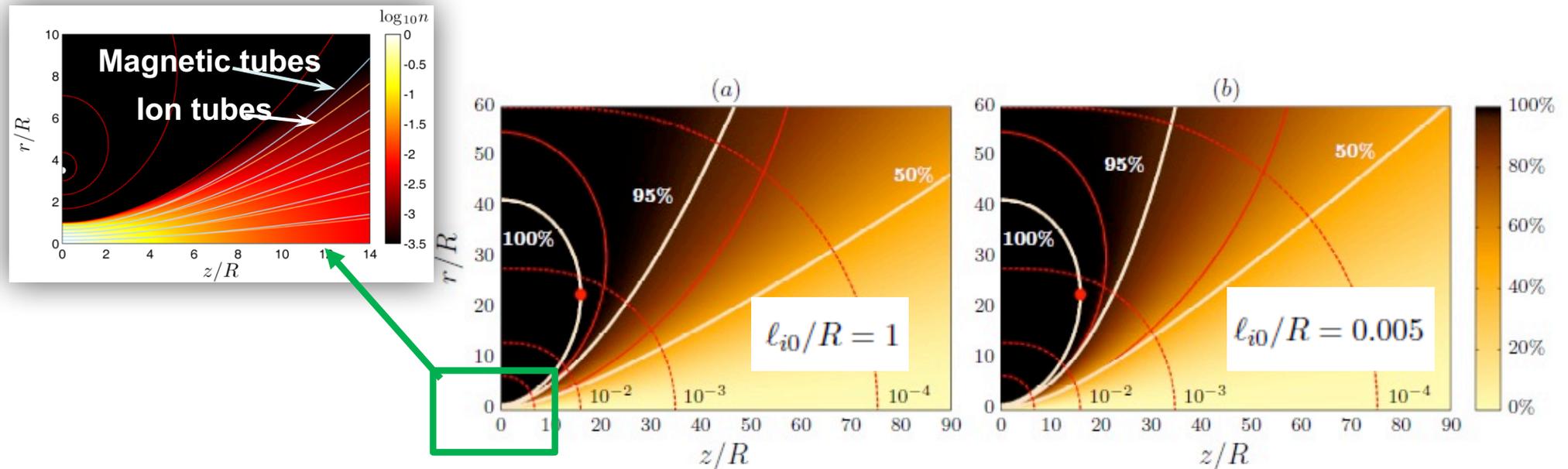
$$m_i n (\mathbf{u}_i \cdot \nabla) \mathbf{u}_i = -en \nabla \phi + en \mathbf{u}_i \times \mathbf{B},$$

- Ions are weakly magnetized in most of the MN
- The ambipolar electric field is self-built in order to preserve quasineutrality (but not current ambipolarity, $\mathbf{j} = \mathbf{0}$).
 - This E-field is not strong enough to make i-tubes = B-tubes



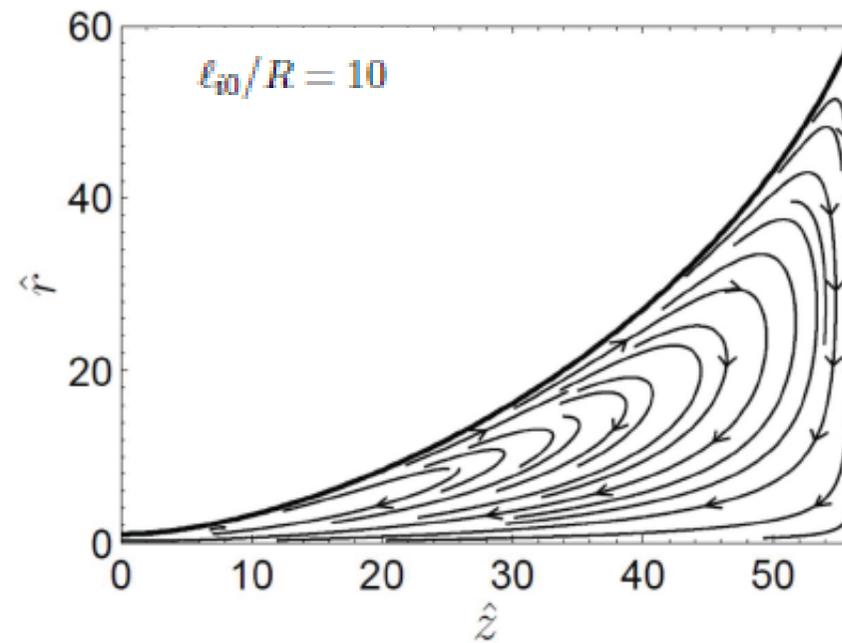
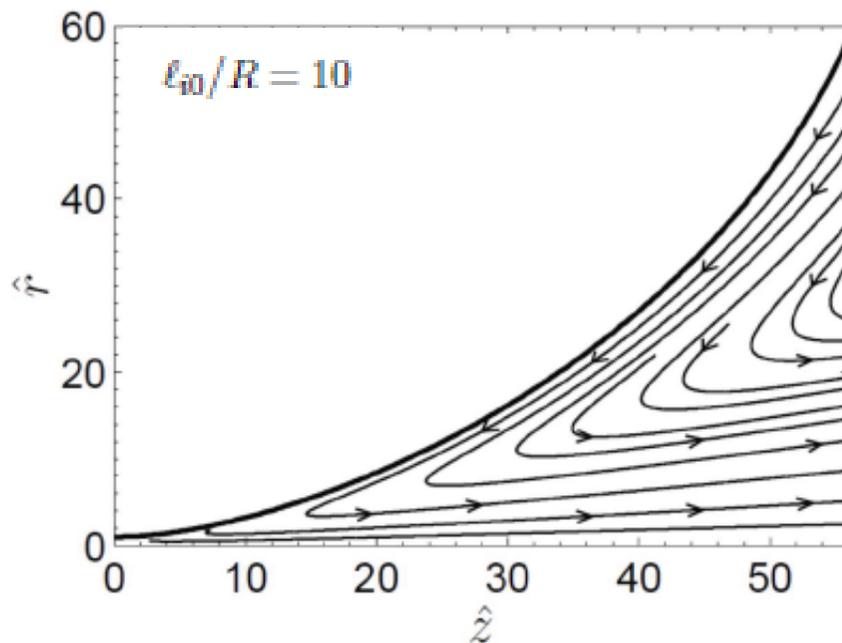
Plasma detachment

- Ion separation increases dramatically after turning point of plasma jet
- When ions become hypersonic, the i-tubes become conical.
- Very small amount of plasma momentum turns back toward the thruster.
- Detachment is mildly sensitive to initial ion magnetization
- In DIMAGNO, electrons still remain perfectly attach to B-lines, but this is not very relevant since they constitute mainly a neutralizing cloud.



Formation of longitudinal loops

- The separation of ion and electron streamtubes produce necessarily longitudinal electric currents, i.e. $\tilde{j} \neq 0$
- How these longitudinal current loops close on themselves, is out of reach of DIMAGNO: coupling with both the source model and beam far-downstream behavior are needed

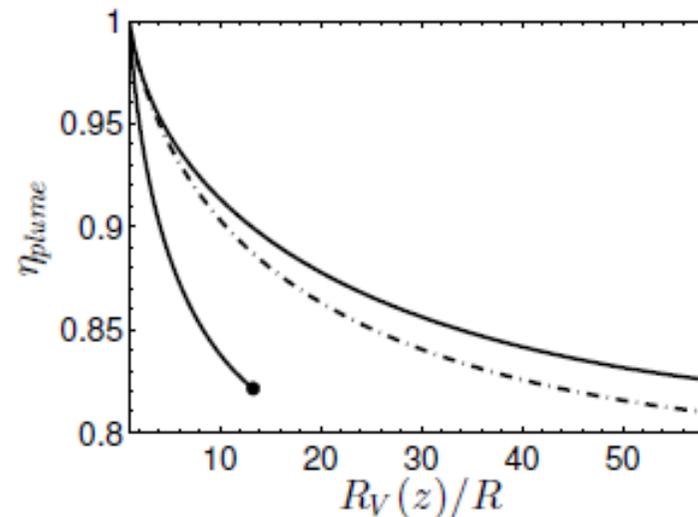


Plume divergence

- DIMAGNO analysis has shown that plasma/MN detachment issue is reduced to an assessment of plume divergence, i.e. how much of the momentum gained in the MN is lost radially.
- Plume divergence efficiency function at B=const section is defined as

$$\eta_{plume}(B) = \frac{\int_{S(B)} dA n u_{zi}^3}{\int_{S(B)} dA n u_i^2 u_{zi}}$$

- Better behavior is obtained for
 - low ion magnetization
 - low MN divergence rate



Increment of plasma momentum along MN

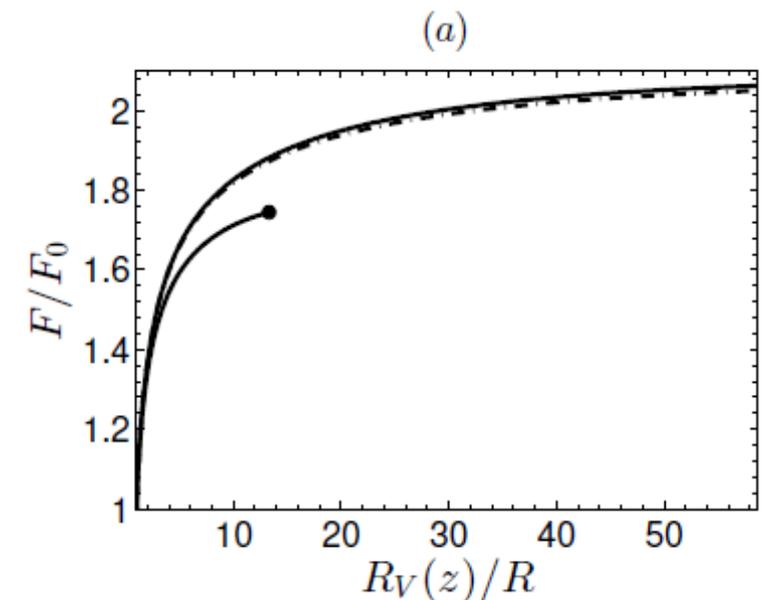
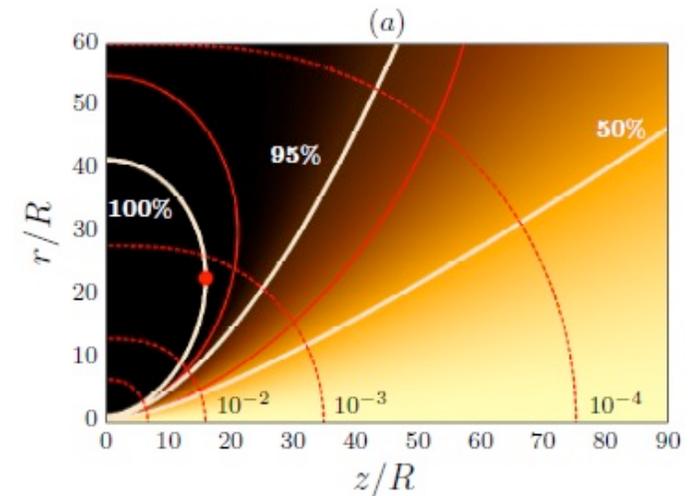
- Conversion of internal-into-kinetic plasma energy (or thermal-into-dynamic plasma momentum) does not produce necessarily thrust:
 - double layers are an example
- Plasma momentum equation (once the intermediary E-field is compensated):

$$\nabla \cdot (m_i n u_i u_i + p_e \bar{\bar{I}}) = j \times B.$$

- Increase of axial plasma momentum is due to magnetic axial force. At B=const sections:

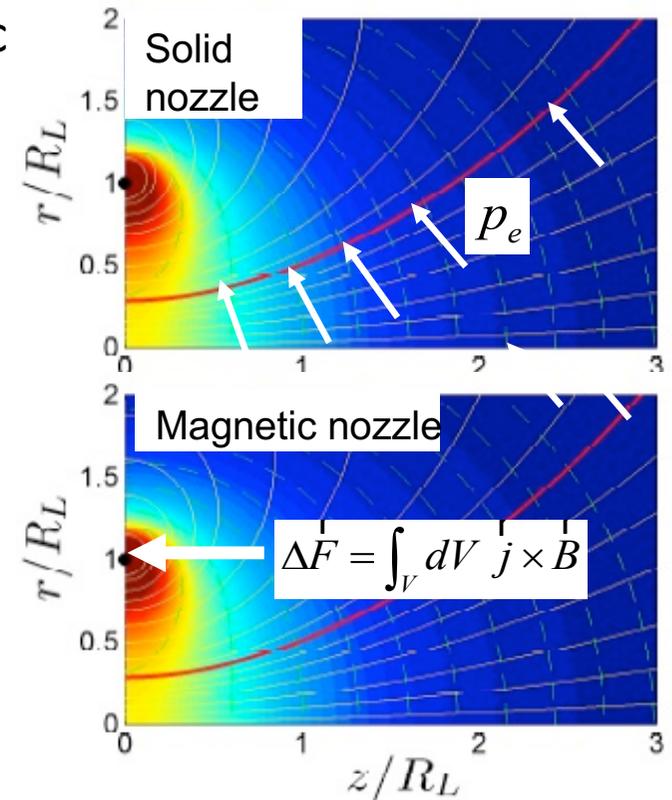
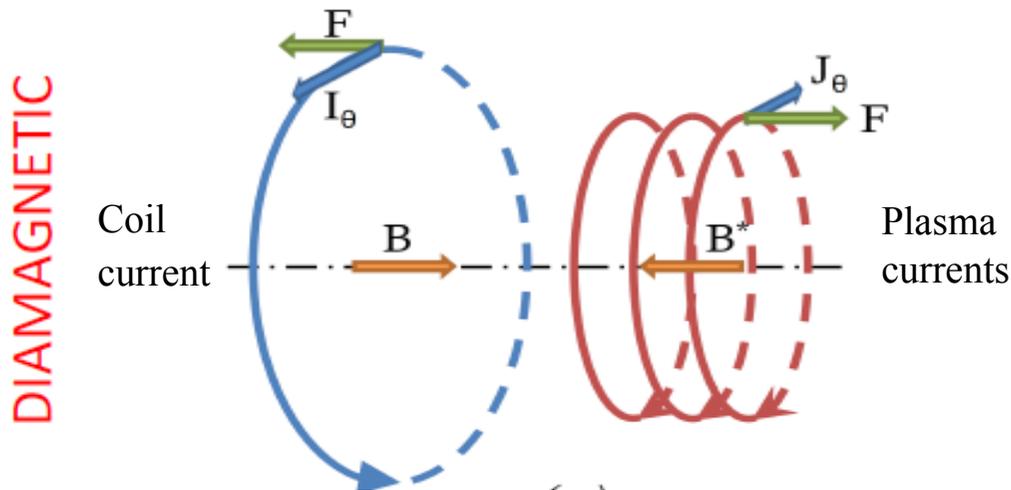
$$F(B) = F_0 + \int_{V(B)} dV (-j_\theta) B_r,$$

with F_0 =axial momentum at the exit of the plasma source (mostly from pressure on rear wall)



Azimuthal currents & Magnetic thrust

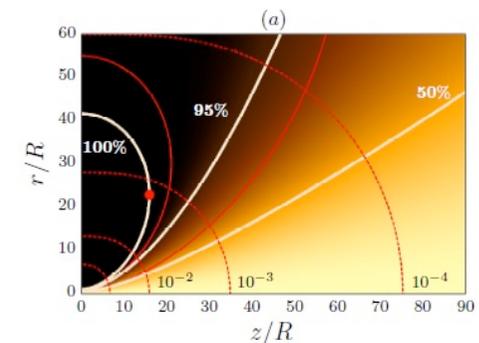
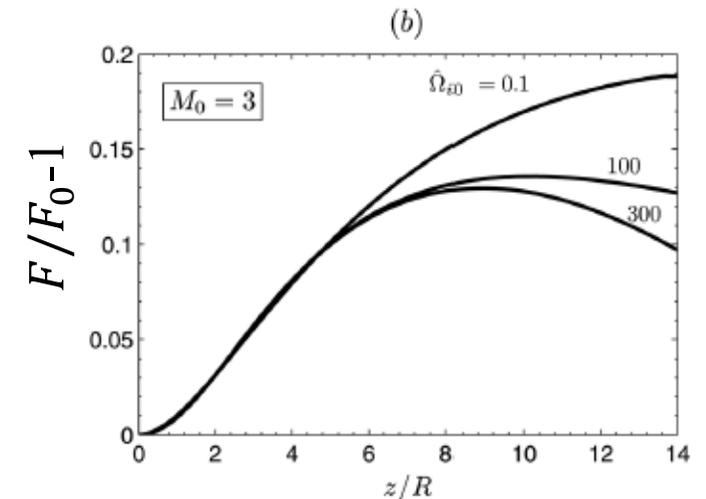
- Magnetic thrust mechanism based on 3rd Newton's law:
 - (Large) B-field from coils \Rightarrow axial force on plasma (travelling, small) θ -current loops
 - (Small) B-field from plasma \Rightarrow axial force on coil (large) θ -currents
- Currents must counterflow for positive plasma acceleration & thrust
- Ion azimuthal currents are small but paramagnetic and detrimental (they produce drag)



On ion & electron dynamics

- 1) MNs are inefficient for a supersonic beam, since thrust/Isp gain is marginal...and can even be negative
- 2) The dynamics of ions and electrons are very different (even leaving apart their different magnetization level).
 - Ions are accelerated almost freely by the E-field.
 - Ion fluid and individual ions behave similarly
 - Individual and fluid electron responses are different.
 - Individual electrons are confined by the E-field and most of them bounce back-and-forth along \mathbf{B}
 - Electron fluid velocity along \mathbf{B} is just the contribution of the small fraction of electrons escaping downstream with ions.
 - Azimuthal e-motion combines $E \times B$ (particle) and $\nabla p_e \times B$ (fluid)

$$u_i(\infty, 0) = c_{s0} \sqrt{M_0^2 + \frac{2}{\gamma - 1}},$$



FUMAGNO: Full ion magnetization model (I)

- Zero ion Larmor radius is a degenerate limit of 2D DIMAGNO
 - Both ions and electrons remain tied to B-streamtubes

- Ion differential equations transform into conservation relations along the magnetic tubes:

$$u_{\perp i} = 0,$$

$$u_{\theta i} \rightarrow 0,$$

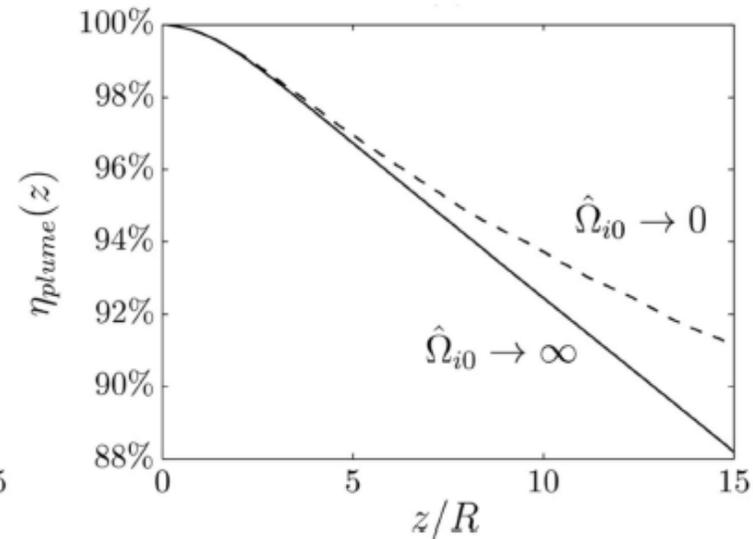
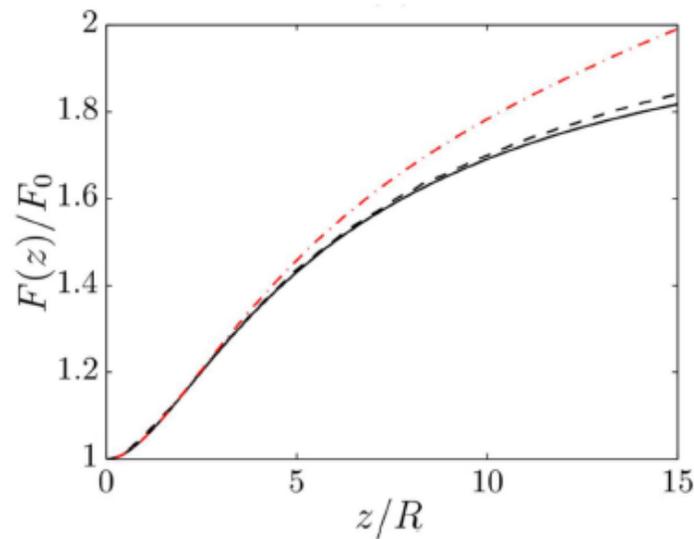
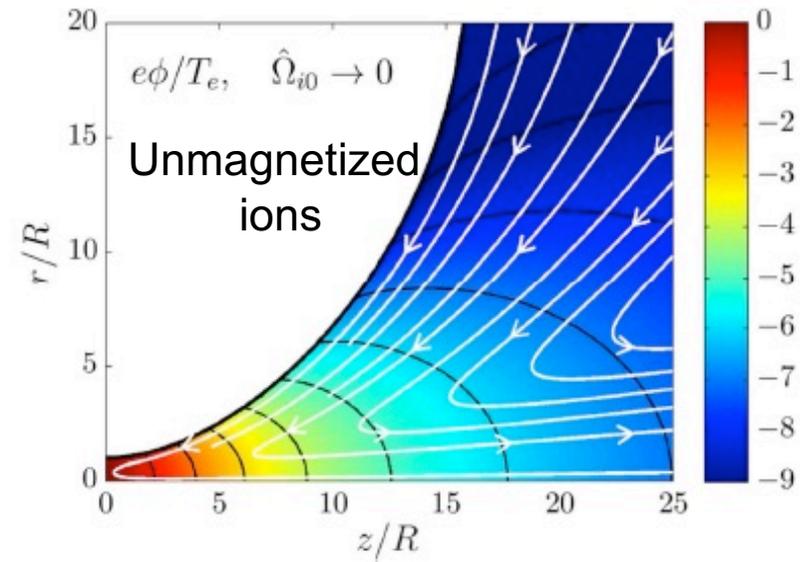
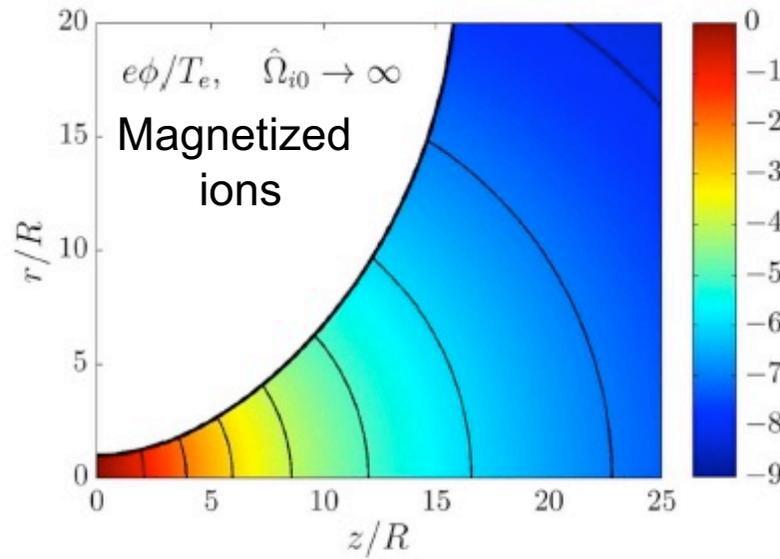
$$eu_{\theta i}B = -eE_{\perp} + \kappa_B m_i u_{\parallel i}^2,$$

$$m_i u_{\parallel i}^2 / 2 + e\phi = H_i(\psi),$$

$$u_{\parallel e} = u_{\parallel i},$$

- Notice the asymptotic behavior of the azimuthal velocity and force
- FUMAGNO solution is practically analytical in each B-streamline
 - Clear improvement over other published approximate models
- FUMAGNO yields a good approximation of MN near region and of magnetic thrust
- FUMAGNO is not applicable at all in the far region and to study plasma separation/detachment

FUMAGNO (II)



Beyond DIMAGNO & FUMAGNO

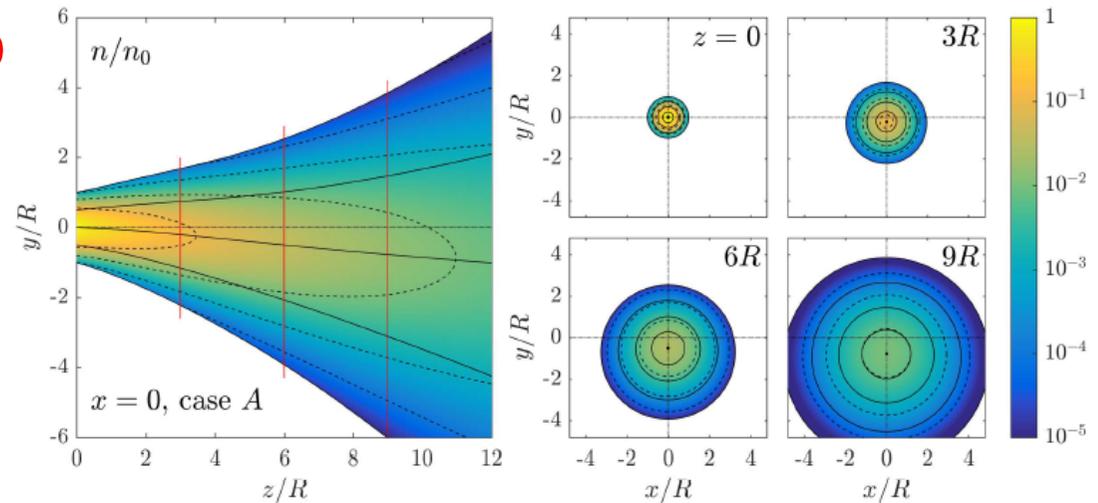
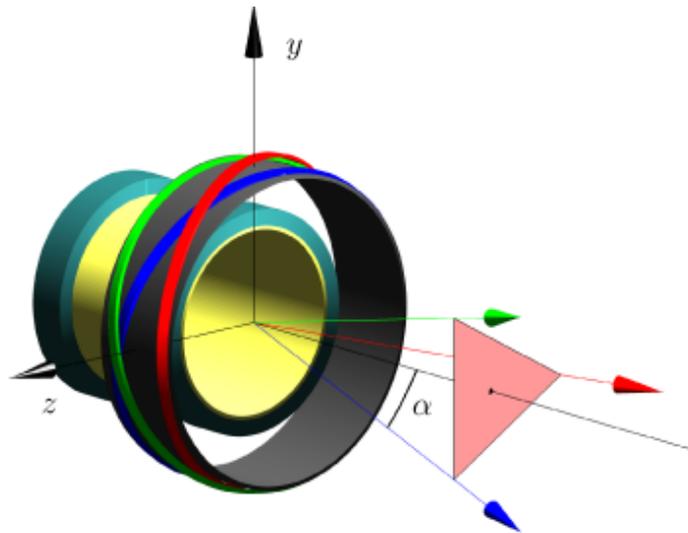
- DIMAGNO assumptions are being relaxed/modified in order to study more physics and understand more phenomena
- **Main open problems:**
 - Energetic ions
 - Electron collisionality
 - 3D magnetic nozzle
 - Full coupling with plasma source
 - Induced magnetic field (*)
 - Electron demagnetization (*)
 - Collisionless electron cooling (*)
 - Instabilities , anomalous diffusion?

(*) These effects are important to close properly the MN very-far region

VECMAN: 3D magnetic nozzle

- Thrust vector control mechanism with no movable parts (patented)
 - Set of several coils placed in oblique planes.
- 3D MN has been studied with 3D version of FUMAGNO
 - Good for near region

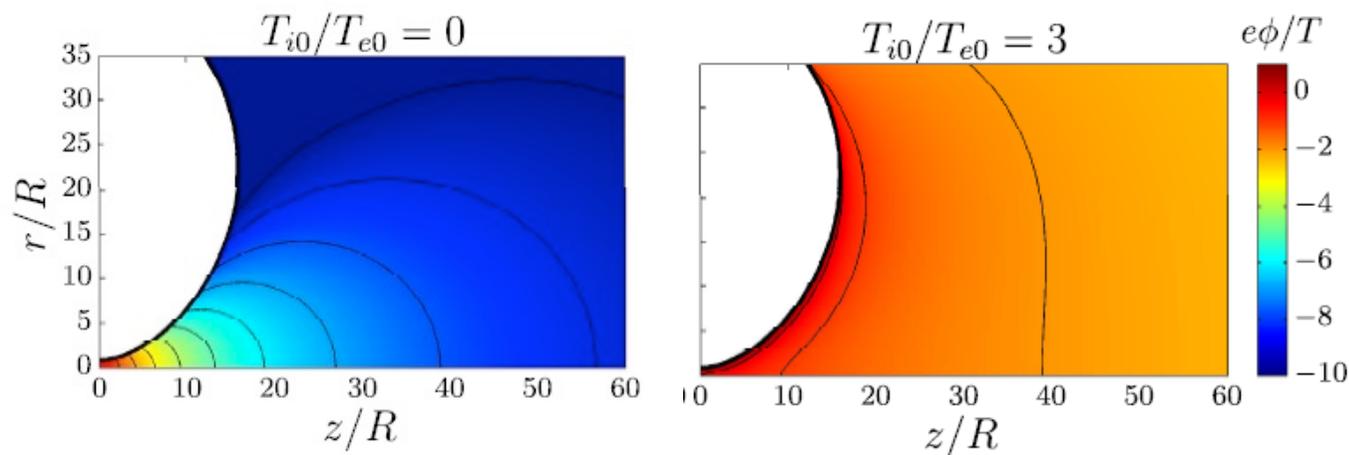
- Pending problem: 3D-DIMAGNO
Important issues related to:
electron & ion confinement,
different integration scheme



Simulation	Ampere-turn ratios	F/F_0	ψ (deg)	θ (deg)	θ_B (deg)
<i>O</i>	15 : 0.33 : 0.33 : 0.33	1.44	—	0.00	0.00
<i>A</i>	15 : 1 : 0 : 0	1.44	-180.00	5.66	5.76
<i>B</i>	15 : 0.5 : 0.5 : 0	1.44	-120.00	2.86	2.91
<i>A'</i>	15 : 5 : 0 : 0	1.34	-180.00	11.06	11.24
<i>B'</i>	15 : 2.5 : 2.5 : 0	1.34	-120.00	5.61	5.70

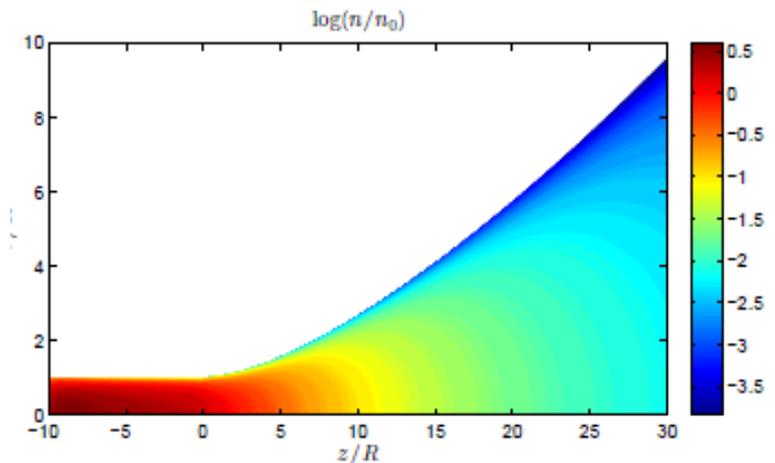
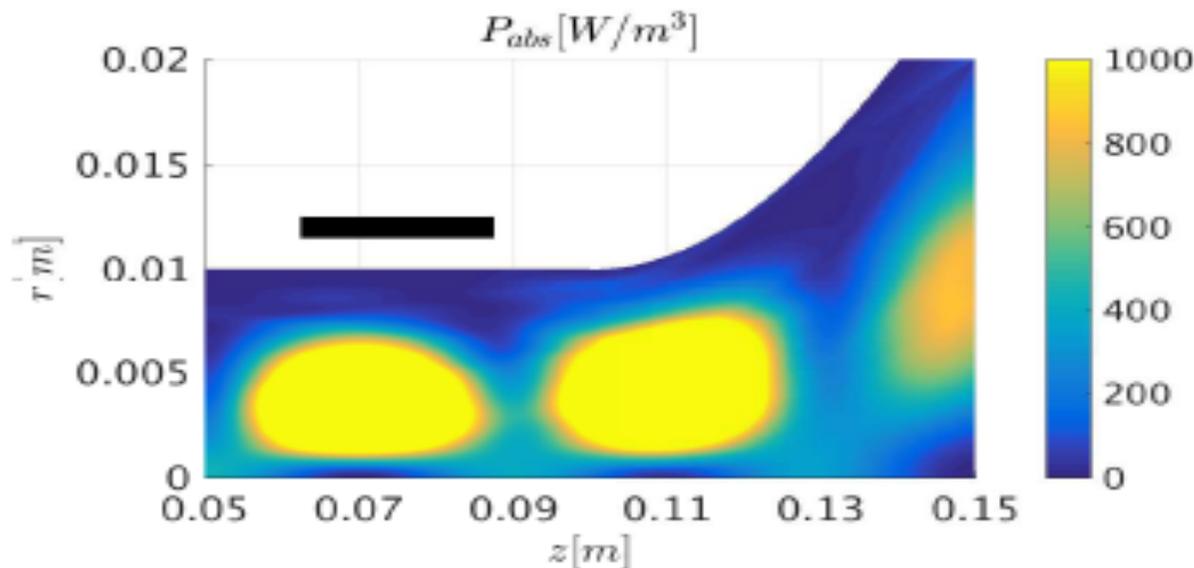
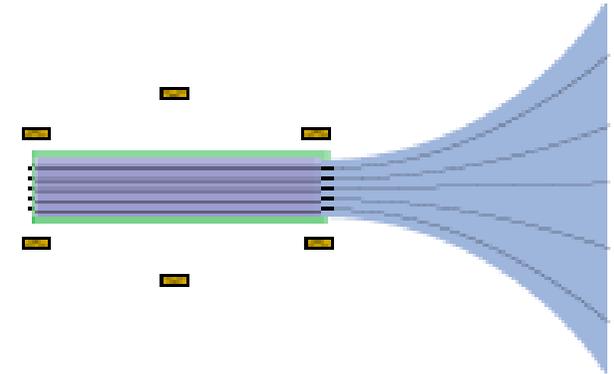
Energetic ions. Several cases

- Source producing hot ions with $T_i \geq O(T_e)$. Included in DIMAGNO.
 - Uncertainties on throat conds. (coming from those on source model)
 - Ions acquire kinetic energy from both: (1) fluid-dynamic conversion of i-thermal energy; (2) electrothermal conversion of e-thermal energy
 - Electric potential fall depends only on T_e .
- **(AF-MPDT)** Ions acquire swirl (kinetic) energy, $u_{\theta i}$.
 - An AF-MPDT **source model is needed** before implementation in MN
- **(VASIMR)** Ions are energized perpendicularly to B and coherently.
 - MN ion equations need to be modified. **Source model is needed** too.



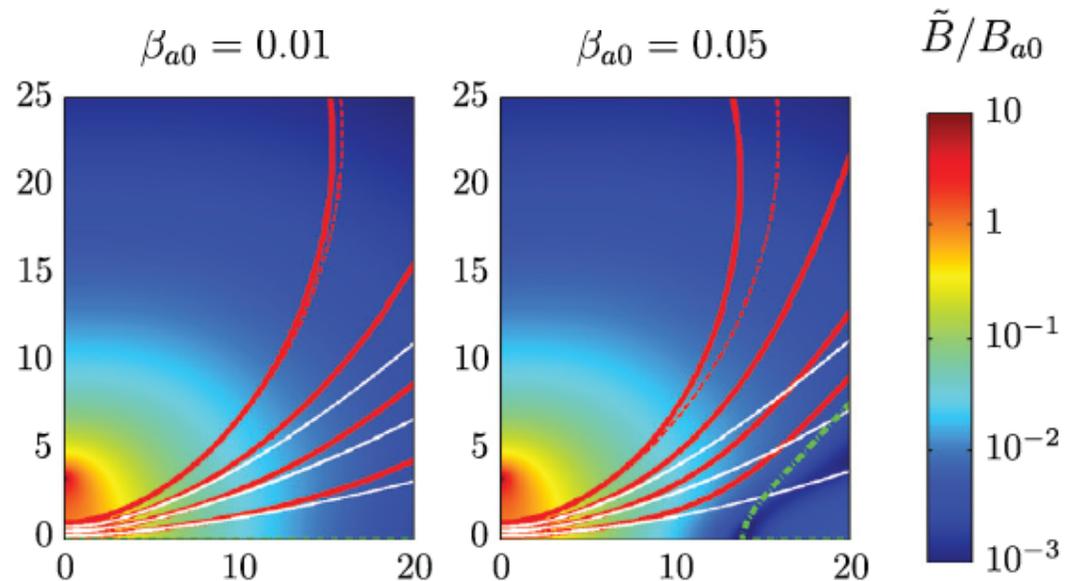
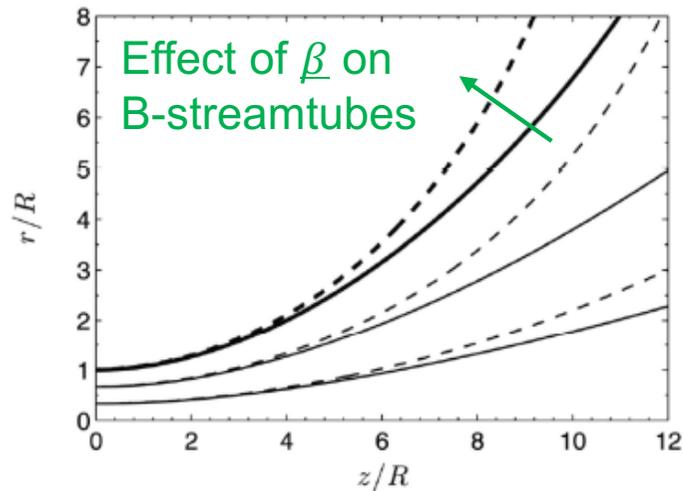
Helicon source + MN = HPT

- Coupling of MN with plasma source not fully solved
 - Fluid transport
 - Wave-plasma coupling
- Energy transfer and T_e field are key aspects
- Numerical algorithms are very sensitive in MN (short wavelenths, oblique B-field)



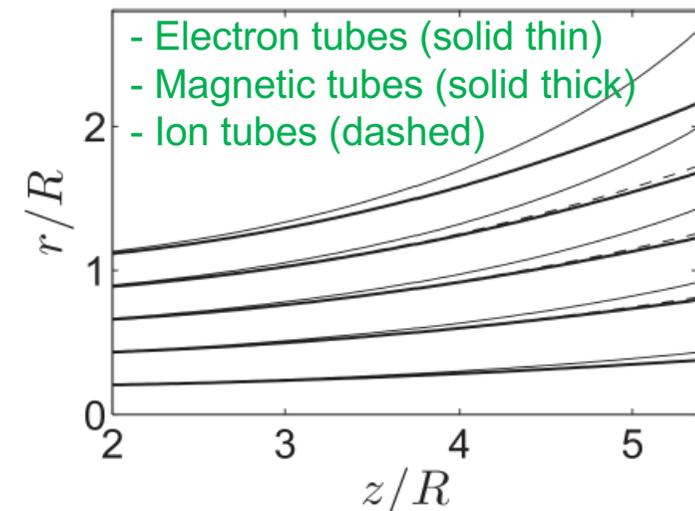
Induced magnetic field

- Azimuthal plasma currents induce a longitudinal B-field $\nabla \times \tilde{\mathbf{B}}^* = \mu_0 j_\theta \mathbf{1}_\theta$.
- We saw that this field is crucial for thrust transmission
- It affects MN shape and plasma expansion only if $\beta = \frac{\mu_0 n T_e}{B^2} = O(1)$
- The induced field opposes the applied one \Rightarrow total B-field is weaker, \Rightarrow MN and plume divergences increase
- Open issues: (1) problem is elliptic, current loops affect upstream region
(2) islands of zero B-field (no-magnetization) are created



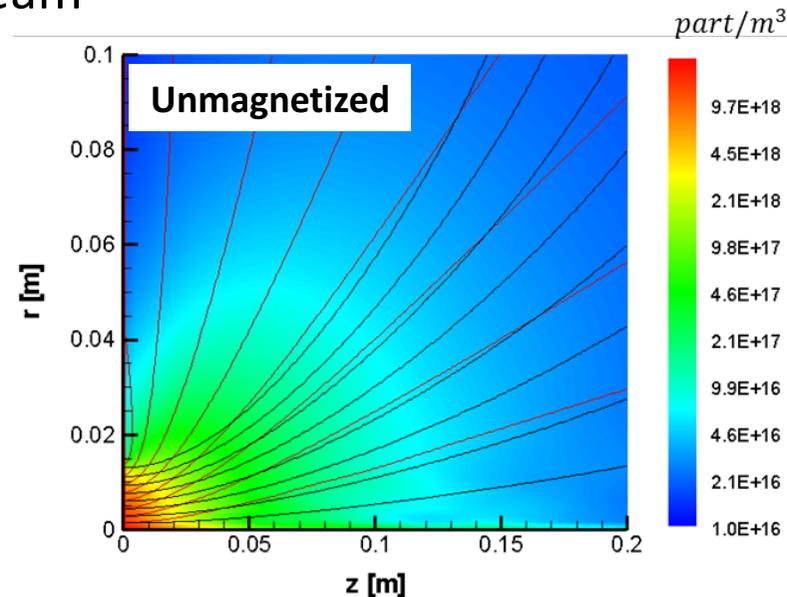
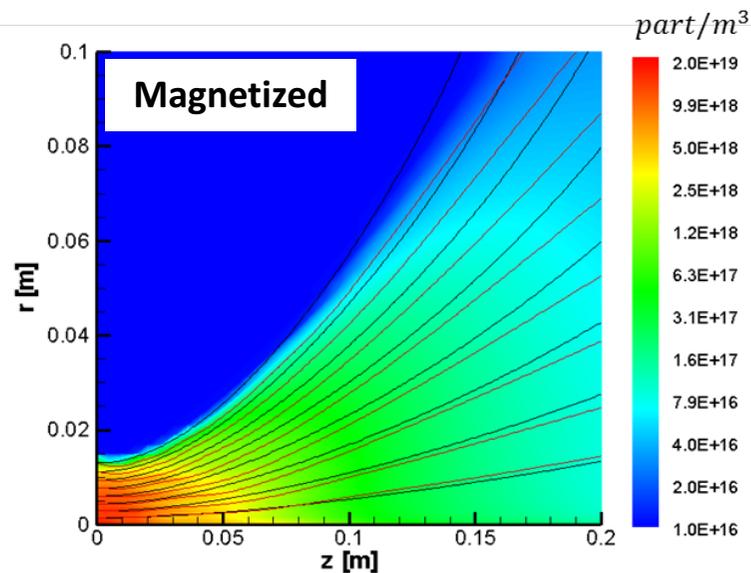
Electron demagnetization

- The transition (electron) magnetized-to-unmagnetized plume is crucial for the downstream plasma beam closure
- This is to us the main open problem in our MN studies
- DIMAGNO operates under the zero e-inertia, zero e-gyroradius limits.
- Nonzero e-gyroradius effects appear as B decreases.
 - These effects (which include e-inertia and pressure anisotropy) are complex to model consistently
 - Some advances have been made extending DIMAGNO
 - Electron demagnetization tends to increase plasma beam divergence
 - Furthermore, a perturbation approach is possibly non worthy, since e⁻ become demagnetized further downstream.
- Key point: How to reconcile magnetized and unmagnetized e-models



Electron collisionality

- The key parameter is the (finite) Hall parameter $\chi = \Omega_e/\nu_e$
- Collisionality (with neutrals and ions, or anomalous) separate electrons from magnetic streamtubes, which diffuse outwards with $u_{\perp e} = \frac{\nu_e}{\Omega_e} u_{\theta e}$
- Collisional effects increase downstream



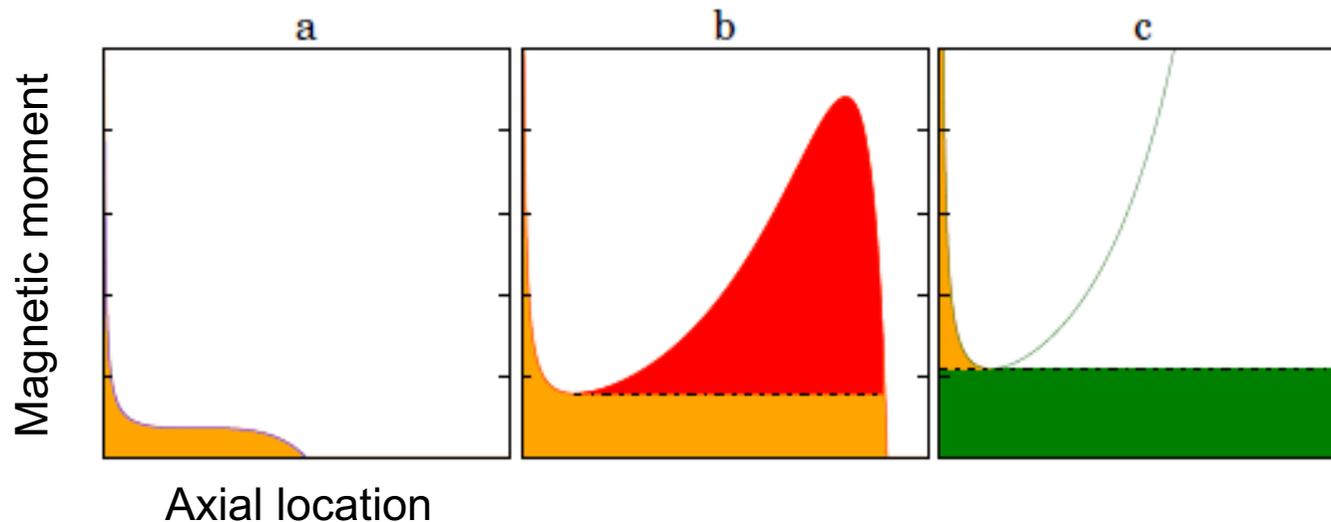
- These plots were obtained with a fluid/PIC MN code under development
- Open issues: (1) effect CEX collisions (2) far region solution in order to verify that MN prevents most of ion backflow

Collisionless electron cooling (I)

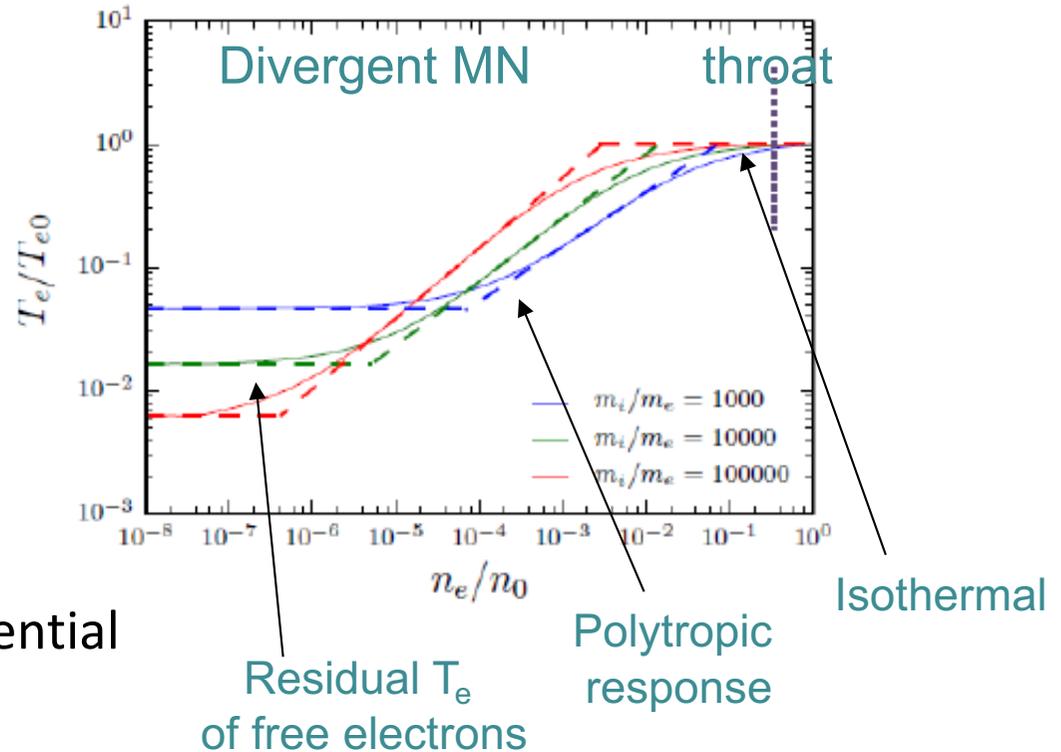
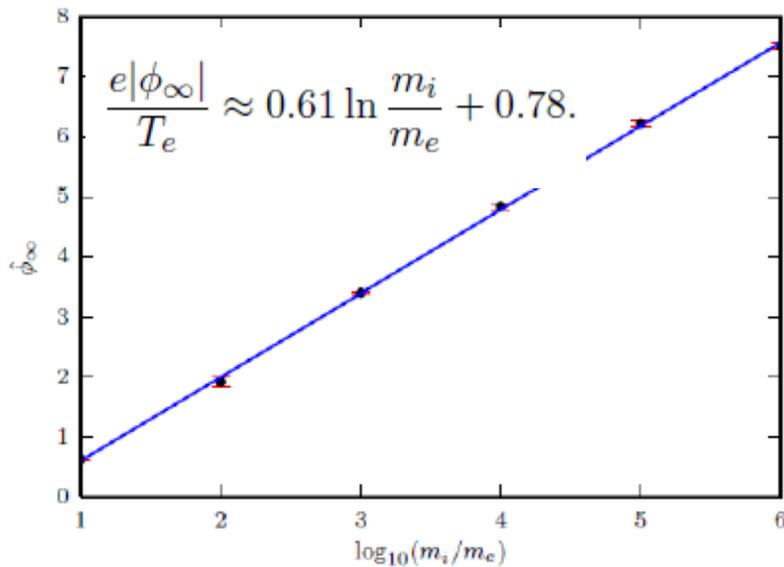
- Plasma beam is rarefied \Rightarrow few collisions \Rightarrow no local thermodynamic equilibrium \Rightarrow no justification for isentropic/adiabatic behavior
- Isothermal behavior not applicable to expansion in an infinite region
- **Kinetic treatment of electron population is required**
(Much less important for ions)
- This open problem is common to magnetized and unmagnetized plumes
 - It could be easier to treat in magnetized plumes
- Experience shows that there is some electron cooling
 - Phenomenological polytropic fittings are used
- Amount of plume cooling is going to determine final ion energy and interaction with SC surfaces \rightarrow **critical issue for SC operators**
- Advances have been made with a kinetic model of paraxial convergent-divergent MN with a collisionless, fully magnetized plasma
 - Particles conserve their total energy and their magnetic moment

Collisionless electron cooling (II)

- Electrons from upstream source are reflected back when $v_{\text{axial}} = 0$ (*line*)
- Different regions in the EVDF space
 - Void region (*white*) → Main responsible of cooling
 - Region of free electrons (*green*)
 - Region of reflected electrons (*yellow*)
 - Islands of doubly-trapped electrons (*red*) → Origin in collisions or MN formation, Very important in solution → **How to characterize them?**
- **Beyond the paraxial case, problem is totally unexplored**



Collisionless electron cooling (III)



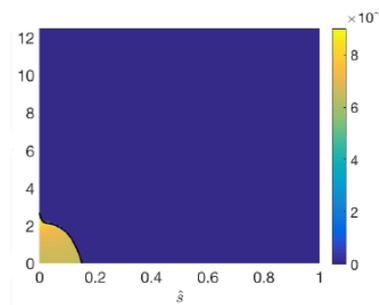
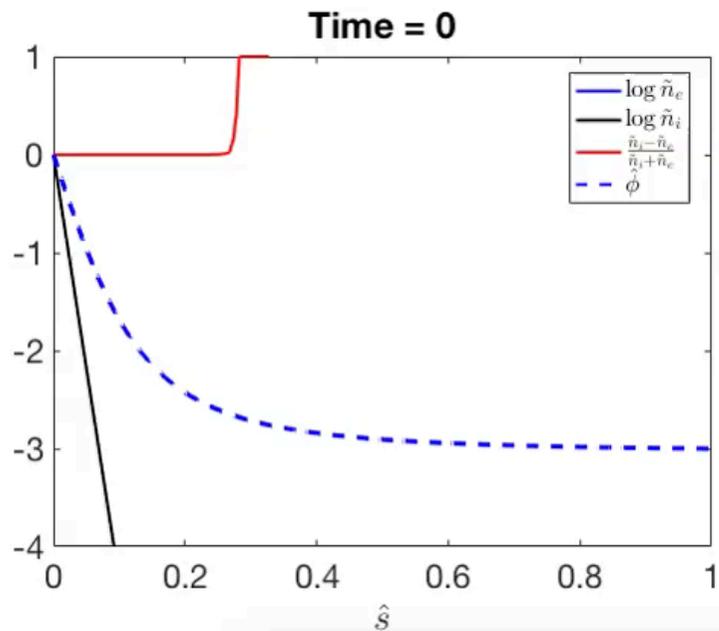
- Interesting comparison with potential fall in Debye sheath:

$$\frac{e\phi_{sh}}{T_e} \approx 0.5 \ln \frac{m_i}{m_e} - 0.92$$

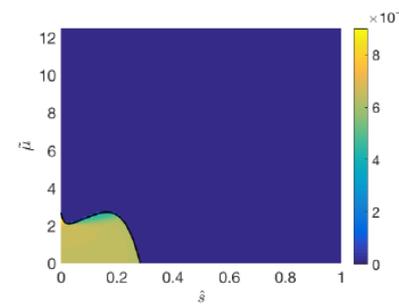
- Main differences for potential fall in MN:
 - It develops in an infinite region not in a very thin region
 - It is fully quasineutral
 - EVDF is undetermined and much more complex

Collisionless electron cooling (IV)

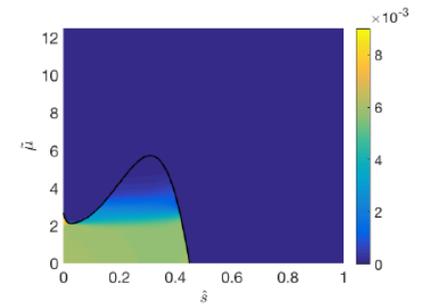
- A time-dependent paraxial kinetic MN model is being developed to characterize the filling of doubly-trapped electron regions
 - Only partial filling is accomplished
- **Open issue: Will occasional collisions eventually fill totally these regions?**



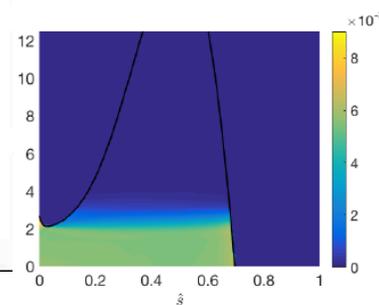
(a) $\tilde{t} = 120$



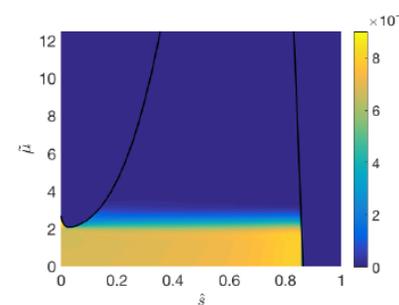
(b) $\tilde{t} = 240$



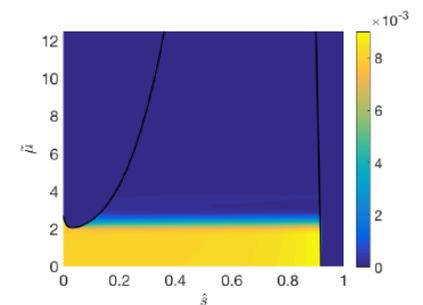
(c) $\tilde{t} = 360$



(d) $\tilde{t} = 540$



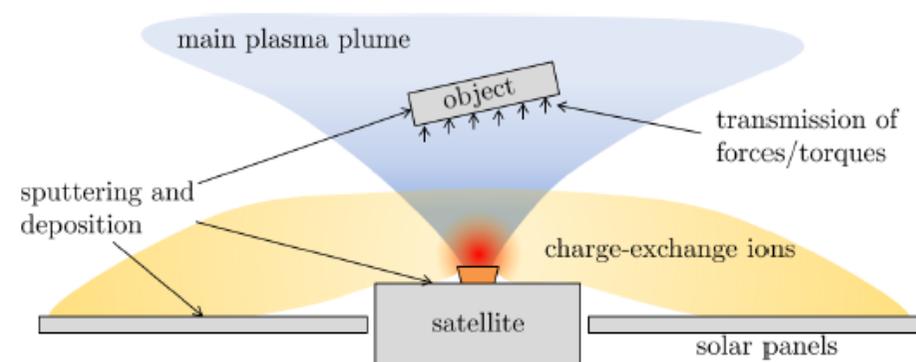
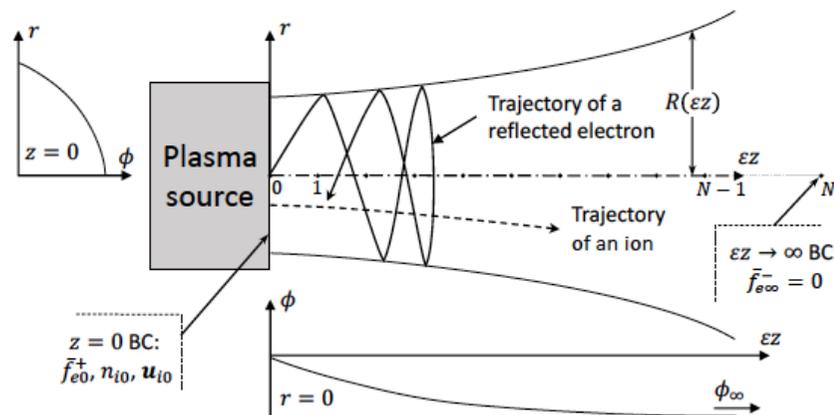
(e) $\tilde{t} = 840$



(f) $\tilde{t} = 2400$

Collisionless electron cooling (V)

- Electron cooling takes place in unmagnetized plume too
 - Now electron motion is strongly 2D (not circumscribed to a B-line)
- Since a radial electron bouncing is confined and fast:
 - The adiabatic invariant J_r (action integral in r orbits) exists playing a similar role to magnetic moment μ in magnetized plumes
- This model is near finalization and characterizes cooling dependence on plume properties (ion Mach number, propellant type, T_{e0})
- **Open problem: cooling in high-angle directions. Critical for SC interaction**



Our paper publications on MN

1. E. Ahedo & M. Merino, "Two-dimensional supersonic plasma acceleration in a magnetic nozzle", Physics of Plasmas 17, 073501(2010) [[PDF](#)] [[DOI](#)]
2. E. Ahedo & M. Merino, "On plasma detachment in propulsive magnetic nozzles", Physics of Plasmas 18, 053504 (2011) [[PDF](#)] [[DOI](#)]
3. M. Merino & E. Ahedo, "Simulation of plasma flows in divergent magnetic nozzles", IEEE Transactions on Plasma Science 39, 2938-2939 (2011) [[PDF](#)] [[DOI](#)]
4. E. Ahedo & M. Merino, "Two-dimensional plasma expansion in a magnetic nozzle: separation due to electron inertia", Physics of Plasmas 19, 083501 (2012) [[PDF](#)] [[DOI](#)]
5. M. Merino & E. Ahedo, "Two-dimensional quasi-double-layers in two-electron-temperature, current-free plasmas", Physics of Plasmas 20, 023502 (2013) [[PDF](#)] [[DOI](#)]
6. M. Merino & E. Ahedo, "Plasma detachment in a propulsive magnetic nozzle via ion demagnetization", Plasma Sources Science and Technology 23, 032001 (2014) [[PDF](#)] [[DOI](#)]
7. M. Merino & E. Ahedo, "Influence of Electron and Ion Thermodynamics on the Magnetic Nozzle Plasma Expansion", IEEE Transactions on Plasma Science 43, 244-251 (2015) [[PDF](#)] [[DOI](#)]
8. M. Merino & E. Ahedo, "Fully magnetized plasma flow in a magnetic nozzle", Physics of Plasmas 23, 023506 (2016) [[PDF](#)] [[DOI](#)]
9. M. Merino & E. Ahedo, "Effect of the plasma-induced magnetic field on a magnetic nozzle", Plasma Sources Science and Technology 25, 045012 (2016) [[PDF](#)] [[DOI](#)]
10. M. Merino & E. Ahedo, "Magnetic Nozzles for Space Plasma Thrusters", published in Encyclopedia of Plasma Technology 2, edited by J. Leon Shohet (2016, Taylor and Francis) [[PDF](#)]
11. M. Merino & E. Ahedo, "Contactless steering of a plasma jet with a 3D magnetic nozzle", Plasma Sources Science and Technology , 045012(in press)

Thank you!

Acknowledgments: To EP2 researchers contributing to this topic

uc3m

